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Key Points:

- The first multiparameter parameterizations of mixing efficiency and turbulent Prandtl number are proposed
- Global maps of mixing efficiency suggest local estimates that may be grossly different from the canonical value of 0.2
- Estimates of diapycnal diffusivity need to be reassessed based on the generalized-Osborn formula

Supporting Information:

- Supporting Information S1
- Figure S1
- Figure S2

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A new characterization of the turbulent diapycnal diffusivities of mass and momentum in the ocean

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**Abstract** The diapycnal diffusivity of mass supported by turbulent events in the ocean interior plays a fundamental role in controlling the global overturning circulation. The conventional representation of this diffusivity, due to Osborn (1980), assumes a constant mixing efficiency. We replace this methodology by a generalized-Osborn formula which involves a mixing efficiency that varies nonmonotonically with at least two nondimensional variables. Using these two variables, we propose dynamic parameterizations for mixing efficiency and turbulent Prandtl number (the latter quantifies the ratio of momentum to mass diapycnal diffusivities) based on the first synthesis of an extensive direct numerical simulation of inhomogeneously stratified shear-induced turbulence. Data from Argo floats are employed to demonstrate the extent of the spatial and statistical variability to be expected in both the diapycnal diffusivities of mass and momentum. We therefore suggest that previous estimates of these important characteristics of the global ocean require reconsideration.

1. Introduction

The return of abyssal waters to the ocean surface depends critically on the irreversible process of diapycnal mixing [Munk, 1966; Talley, 2013]. However, debate continues concerning the spatial and temporal variability of the efficiency of such mixing in the global oceans [Peltier and Caulfield, 2003; Wunsch and Ferrari, 2004; Ivey et al., 2008]. An equally pressing issue that has not been sufficiently addressed, concerns the degree of momentum diffusion associated with the turbulent diapycnal diffusion of mass. In order to assess the possible spatial variability of these quantities, improved understanding of small-scale stratified turbulent mixing processes must be framed in a global context. Our goal in this paper is to take a significant step toward accomplishing this task by integrating the extensive direct numerical simulation (DNS) data set of Salehipour and Peltier [2015] (referred to hereinafter as [SP]) with the global strain measurements from Argo floats as reported in Whalen et al. [2012, 2015].

The conventional method for relating estimates of turbulent dissipation,  $\epsilon$ , to turbulent diapycnal diffusivity,  $K_p$ , is based on a simplified turbulent kinetic energy (TKE) formula first proposed by Osborn [1980], namely:

$$K_p = \gamma \frac{\epsilon}{N^2} \tag{1}$$

in which  $N^2 = -g/\rho_0(d\bar{\rho}/dz)$ , where  $\bar{\rho}$  denotes the mean flow density, and  $\gamma = R_f/(1-R_f)$  is a “flux coefficient” based on the flux Richardson number,  $R_f$  (defined as the ratio of the buoyancy flux to the shear production of TKE), which is conventionally assumed to be a measure of mixing efficiency. Following the empirical suggestions of Osborn [1980] that  $R_f \leq 0.17$ ,  $\gamma$  is frequently assumed to be fixed at  $\gamma = 0.2$ . We will hereinafter refer to this practice as the Osborn method. This formula suffers from two fundamental problems: (i) its underlying assumptions include homogeneous and statistically stationary flow conditions which are not strictly satisfied by actual oceanographic turbulent flows and (ii) the assumption of constant mixing efficiency is an oversimplification [Smyth et al., 2001; Ivey et al., 2008; Mashayek and Peltier, 2013].

The first problem has recently been resolved in SP in which the accurate representation of diapycnal diffusivity due to Winters and D’Asaro [1996] has been recast into an “Osborn-like” form (referred to hereinafter as the generalized-Osborn formula). Not only is this new formula valid for any Boussinesq flow (i.e., it relaxes the original limiting assumptions on which the Osborn formula is based) but also distinguishes between irreversible

and reversible processes while closely resembling the original Osborn formula. This generalized-Osborn formula has the following form (c.f. (1)):

$$K_p = \Gamma \frac{\epsilon}{N_*^2} \quad (2)$$

in which the flux coefficient  $\Gamma = E/(1 - E)$  depends on the precise definition of mixing efficiency,  $E = \mathcal{M}/(\mathcal{M} + \epsilon)$  which invokes a time-dependent adiabatic restratification of the density field into a background stably stratified reference state whose buoyancy frequency is denoted by  $N_*$  and is associated with a state of minimum potential energy [Winters *et al.*, 1995; Caulfield and Peltier, 2000]. Through this “sorting” process, the buoyancy flux is essentially decomposed into a “reversible” (stirring) and an “irreversible” (mixing) component. Mixing,  $\mathcal{M}$  (refer to SP for its explicit definition), therefore represents the rate at which available potential energy is dissipated irreversibly, whereas stirring is characterized by the reversible energy exchanges between the kinetic and potential energy reservoirs. This important distinction, which is neglected in the Osborn formula (1), is the basis on which the generalized-Osborn formula (2) is derived.

As to the second problem, the actual variability of mixing efficiency has remained enigmatic despite the consensus that it cannot be constant [Ivey *et al.*, 2008]. The current parameterizations of mixing efficiency (e.g., that of Shih *et al.* [2005]) are also highly contentious since all of them depend on a single parameter and are therefore inherently ambiguous [Mater and Venayagamoorthy, 2014]. Consequently, our aim is to propose the first multiparameter representation for  $E$ .

Besides the scalar diapycnal diffusivity, all numerical models that do not resolve subgrid scale turbulent processes also rely on a turbulence closure model that prescribes parameterizations for the momentum diffusivity ( $K_m$ ) [e.g., Klymak *et al.*, 2010]. It is a common practice to use the turbulent Prandtl number for this purpose in order to relate the momentum and scalar diffusivities as  $Pr_t = K_m/K_p$ . To properly distinguish between reversible and irreversible contributions to the momentum flux, SP proposed a new formulation for  $K_m$  as

$$K_m = \frac{1}{1 - E} \left( \frac{\epsilon}{S^2} \right) \quad (3)$$

in which  $S = d\bar{u}/dz$  denotes the vertical shear of streamwise mean flow. Because (3) may be rewritten as  $K_m = (Ri_*/E)K_p$ , where  $Ri_* = N_*^2/S^2$  is the gradient Richardson number based on the sorted density profile, a new formulation for  $Pr_t$  as  $Pr_t = Ri_*/E$  is also implied which significantly extends the homogeneous formulations of Venayagamoorthy and Stretch [2010] to any Boussinesq flow and may be contrasted with its conventional definition as  $Pr_t = Ri/R_f$  (see [SP] for further details).

A value of unity for  $Pr_t$  has commonly been assumed based on a physical argument that in a highly turbulent flow the turbulent diffusion rates of the momentum and scalar fields should be equal. For stably stratified flows, different empirical parameterizations have been proposed based on the gradient Richardson number that suggest  $Pr_t$  increases at higher  $Ri$  with  $Pr_t \approx 1$  at low  $Ri$  [Schumann and Gerz, 1995; Venayagamoorthy and Stretch, 2010]. Again a key question of interest to us concerns the spatial variability of  $Pr_t$  throughout the world's oceans. In particular, we investigate whether the current assumption of a globally constant  $Pr_t = 1$  (or even  $Pr_t = 10$ ) [Danabasoglu *et al.*, 2012] is reasonable.

## 2. Methodology

### 2.1. Toward a Multiparameter Mixing Efficiency Parameterization

The multiparameter dependence of mixing efficiency in stratified shear flows suggests that  $E$  must be prescribed in at least a three-dimensional parameter space, for example, as a function of the molecular Prandtl number ( $Pr = \nu/\kappa$ , where  $\nu$  is the kinematic viscosity and  $\kappa$  is the molecular diffusivity), the gradient Richardson number ( $Ri = N^2/S^2$ ) and the buoyancy Reynolds number ( $Re_b = \epsilon/(\nu N^2)$ ). Recent investigations of the stratified turbulence generated by breaking nonlinear Kelvin-Helmholtz (KH) waves at sufficiently high Reynolds numbers explicitly demonstrate such a dependence on  $Ri$  [Mashayek *et al.*, 2013; SP],  $Pr$  [Salehipour *et al.*, 2015], and  $Re_b$  [SP]. While it is certain that  $E$  depends on  $Pr$  [Smyth *et al.*, 2001; Salehipour *et al.*, 2015], we assume for present purposes that the dependences on  $Re_b$  and  $Ri$  are dominant. This assumption reduces the representation of  $E$  to a two-dimensional parameter space expressed as  $E = f(Ri, Re_b)$ .

In order to infer  $E = f(Ri, Re_b)$ , we employ the DNS data set of SP as well as two additional DNS analyses of the same problem but with extremely high values of the initial Reynolds number at  $Re = 2 \times 10^4$  and  $Re = 3 \times 10^4$  (see the supporting information for further details). Furthermore, we split the parameter space

into a region where  $E$  rises with  $Re_b$  (left flank) and a region in which  $E$  decreases with  $Re_b$  (right flank), respectively, associated with the buoyancy-dominated and shear-dominated regimes [SP]. We denote the peak value of  $E$  at which the rollover occurs and its corresponding  $Re_b$  as  $E^*$  and  $Re_b^*$ , respectively.

Our DNS-constrained representation of  $E(Ri, Re_b)$  may then be written as

$$\frac{E}{E^*} = \frac{(1 + 2p) \left( \frac{Re_b}{Re_b^*} \right)^p}{1 + 2p \left( \frac{Re_b}{Re_b^*} \right)^{p+0.5}} \quad (4)$$

in which  $p = 0.55$  if  $Re_b \leq Re_b^*$  and  $p = 1$  if  $Re_b > Re_b^*$  representing the left and right “flanks” respectively. The  $Ri$  dependence is embedded explicitly in  $E^*$  and  $Re_b^*$ , defined respectively as follows:

$$E^*(Ri) = \frac{3 \left( \frac{Ri}{Ri^*} \right)}{8 + \left( \frac{Ri}{Ri^*} \right)^9}, \quad (5)$$

$$Re_b^*(Ri) = \frac{4}{9} \left( \frac{\Psi}{E^*} \right)^2, \quad (6)$$

$$\Psi(Ri) = 0.04 \exp(12Ri) + 1.5, \quad (7)$$

where  $Ri^* = 0.4$  in (5) denotes the  $Ri$  value associated with the peak value of  $E^*$  in its nonmonotonic variation with  $Ri$ . Furthermore,  $Re_b^*(Ri)$  in (6) essentially employs a supplementary variable,  $\Psi$  (7), that is obtained by setting  $E \sim \Psi Re_b^{-0.5}$  in the limit of  $Re_b/Re_b^* \rightarrow \infty$ .

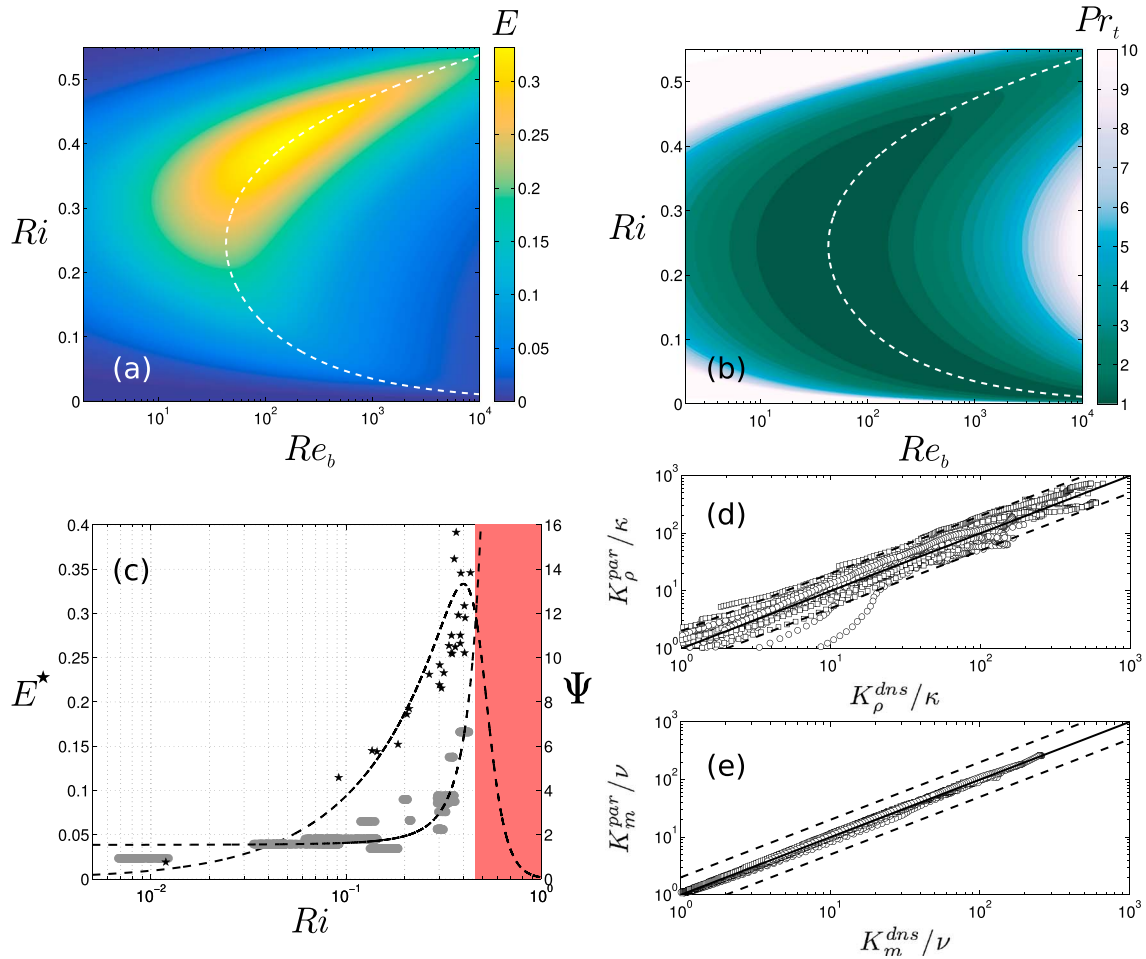
Note that in this paper, both the  $N$  and  $S$  profiles are first squared and then vertically averaged over the depth of the turbulent patch. The resulting values of  $Ri$  inferred from our DNS database hence lie in the range  $0 < Ri < 0.5$  which partly covers the observed values of  $Ri$  in the ocean [see, e.g., Polzin, 1996; Smyth and Moum, 2013; van Haren et al., 2014]. For the purpose of this paper, we have assumed  $E \rightarrow 0$  as  $Ri \rightarrow 1$ .

Figure 1a illustrates the distribution of  $E$  in the  $Ri - Re_b$  space as parameterized by (4). The  $Ri$  variation of  $Re_b^*$  as per (6) is also shown in this figure by a dashed curve. Furthermore, Figure 1c plots  $E^*(Ri)$  and  $\Psi(Ri)$  as obtained from the DNS data together with the inferred nonlinear least squares fits of (5) and (7). The supporting information describes the procedure that was employed to construct the components of this parameterization.

Although it is well established that  $E$  varies nonmonotonically with  $Ri$  [Linden, 1979; Strang and Fernando, 2001; Mashayek et al., 2013; SP], the nonmonotonic dependence of  $E$  on  $Re_b$  [Barry et al., 2001; Shih et al., 2005; SP] has been actively disputed (see, e.g., Gregg et al. [2012] and Bouffard and Boegman [2013] for a discussion). This dual nonmonotonicity has been nonetheless explicitly realized in the DNS analyses of SP and is therefore preserved in the proposed parameterization through the continuous transition between the left and right flanks at  $Re_b^*$  and the nonmonotonic dependence of  $E^*$  on  $Ri$ .

On the “right flank” the  $Re_b^{-0.5}$  power law relation, implicit in (4) with  $p = 1$  when  $Re_b \gg Re_b^*$ , has been adopted here because of the accumulating DNS [Shih et al., 2005; SP] and observational [Davis and Monismith, 2011; Lozovatsky and Fernando, 2013; Bouffard and Boegman, 2013; Bluteau et al., 2013; Walter et al., 2014; Monismith et al., 2015] evidence that mixing efficiency does decrease with  $Re_b$  as  $E = \Psi Re_b^{-0.5}$ . For example, Monismith et al. [2015] reported  $\Psi \sim 1.5 - 4.5$  for a series of observations with  $10^2 < Re_b < 10^7$ . Our DNS data suggest that  $\Psi$  may indeed increase exponentially with  $Ri$  as in (7) and may lie in the range  $1 < \Psi < 15$  for  $0 < Ri < 0.5$ . The high  $Ri$  regions of the parameter space are yet to be explored by DNS of “strongly” stratified turbulence and thus should be interpreted as preliminary. This uncertainty is indicated in Figure 1c by a shaded area.

A number of interesting observations follow from Figure 1a: (i) irreversible mixing becomes as efficient as  $E \sim 1/3$  (or  $\Gamma \sim 0.5$ ) for relatively high levels of stratification with  $Ri \sim 0.35 - 0.45$  which occurs in the range  $\mathcal{O}(10^2) < Re_b < \mathcal{O}(10^3)$ . (ii) This highly efficient mixing is sustained for a broader range of  $Re_b$  if the flow is



**Figure 1.** Contours of (a)  $E = f(Re_b, Ri)$  as parameterized in (4) (to be compared with the fixed canonical value of  $R_f = 0.17$  in the Osborn formula) and (b)  $Pr_t = Ri_*/E$ . The dashed curves in Figures 1a and 1b illustrate  $Re_b^*$  (6). (c) Illustration of  $\Psi(Ri)$  (7) (gray circles) and  $E^*(Ri)$  (5) (black stars) as inferred from DNS analyses. The shaded area in Figure 1c corresponds with the uncertain region of high- $Ri$ . (d, e) Comparison between the parameterized  $K_\rho$  and  $K_m$  (denoted by  $K_\rho^{par}$  and  $K_m^{par}$ ) with their DNS-based origins (denoted by  $K_\rho^{dns}$  and  $K_m^{dns}$ ) at  $Pr = 1$  (circles) and  $Pr > 1$  (squares). The dashed lines identify a factor of 2 underestimation/overestimation.

sufficiently stratified. (iii)  $Re_b^*(Ri)$  varies nonmonotonically with a minimum of  $Re_b^* \sim 40$  at  $Ri = 0.25$  which increases for higher and lower values of  $Ri$  (note that due to the smooth form of  $E$  around  $Re_b^*$ ,  $E$  becomes  $\propto Re_b^{-0.5}$  at  $\sim 9/4Re_b^*$ , e.g., at  $Re_b \sim 90$  for  $Ri = 0.25$ ). The increase in  $Re_b^*$  at higher  $Ri$  occurs because, as the flow becomes more strongly stratified, the transition to the shear-dominated turbulent regime (i.e., right flank) requires increasingly more energetic turbulence and thus  $Re_b^*$  ought to increase. On the other hand,  $Re_b^*$  also has to increase dramatically as  $Ri \rightarrow 0$  since, in order for the resulting buoyancy-dominated turbulence to become shear-dominated, significant turbulent dissipation is required. (iv) The transition to the right flank in *Shih et al.* [2005] at  $Re_b \sim 100$ , as also reinforced by SP, occurs because in both studies high  $Re_b$  was only obtained due to relatively weak stratification. In fact, our exceptionally expensive DNS analyses with initial Reynolds numbers up to  $3 \times 10^4$  have revealed that high  $Re_b$  on the right flank may also be achieved at high  $Ri$  if the Reynolds number is sufficiently high.

To assess the parameterized values of  $E$  against the original DNS-based values, Figure 1d employs the generalized-Osborn formula in (2) to compare the resulting values of  $K_\rho/\kappa$ . As shown in this figure, the DNS data at  $Pr = 1$  are represented quite accurately, while those at higher  $Pr$  are mostly overestimated by the parameterization but are still within a factor of 2. This overestimation is partly because the  $Pr$  effect has been neglected in this preliminary form of the parameterization. Furthermore, the accuracy of the proposed parameterization will be realized by comparing  $K_\rho^{dns}$  against traditional models of *Osborn* [1980] and *Osborn and Cox* [1972] as has been done explicitly in Figure 2 of SP.

## 2.2. Toward a Multiparameter $Pr_t$ Parameterization

The proposed multiparameter representation of  $E$  in (4) directly implies a multiparameter representation of  $Pr_t = g(Ri, Re_b)$  after recalling that  $Pr_t = Ri_*/E$  and  $Ri_* \sim Ri$  (or  $N \sim N_*$ ) due to vertical averaging [SP]. We will hereinafter use  $Ri$  and  $Ri_*$  interchangeably. Notice that in the absence of diapycnal diffusivity (i.e.,  $K_p = 0$ ),  $K_m$  will maintain nonzero values because internal waves would still sustain momentum flux across isopycnals. Therefore, if  $K_p = 0$ ,  $Pr_t$  loses its utility and  $K_m$  is best described as  $K_m = \epsilon/S^2$  (see (3) for  $E = 0$ ).

Figure 1b illustrates such dual dependence in the  $Ri - Re_b$  space (the dashed curve again indicates  $Re_b^*$  (6) and thus distinguishes the left and right flanks). As shown in this figure,  $Pr_t$  increases with  $Ri$  from  $\sim 1$  at relatively low  $Ri$  values in accord with the previous  $Ri$ -based parameterizations. However, in contrast to all previously suggested parameterizations, the dependence on  $Ri$  critically depends also upon  $Re_b$ . For example, at high  $Re_b$  where turbulence is highly energetic,  $Pr_t$  varies nonmonotonically with  $Ri$  and rises to values as high as  $Pr_t \sim \mathcal{O}(10)$  for moderate values of  $Ri$ . This counter-intuitive rise of  $Pr_t$  is due to the right flank decrease of  $E$  with  $Re_b$  at moderate  $Ri$  and clearly is not in accord with the common assumption of  $Pr_t \sim 1$  for intensely turbulent patches. Finally, in a particular region of the  $Ri - Re_b$  space (and not necessarily at low  $Ri$ ) momentum and mass turbulent diffusivities tend to be equal resulting in  $Pr_t \sim 1$ . It is very interesting to note that, this region of the parameter space coincides with that in which mixing is most efficient.

Similar to Figure 1d, Figure 1e evaluates the parameterized momentum diffusivity  $K_m/\nu$  against that obtained directly from the DNS data set based on (3). The excellent agreement should not be surprising (as also noted by SP) because unlike  $K_p$ ,  $K_m$  is almost insensitive to  $E$  unless  $E \rightarrow 1$  (c.f. (2) and (3)).

## 2.3. Application to the Global Oceans

In order to apply the DNS-constrained parameterizations to the global ocean, we will rely on two major assumptions:

1. The mixing properties of the idealized KH-ansatz are statistically representative of individual oceanic mixing events.
2. The *average* diapycnal mixing supported by a fine-scale internal wavefield is similar to that induced by a turbulent patch, *provided that* their corresponding averaged values of  $Re_b$  and  $Ri$  are similar.

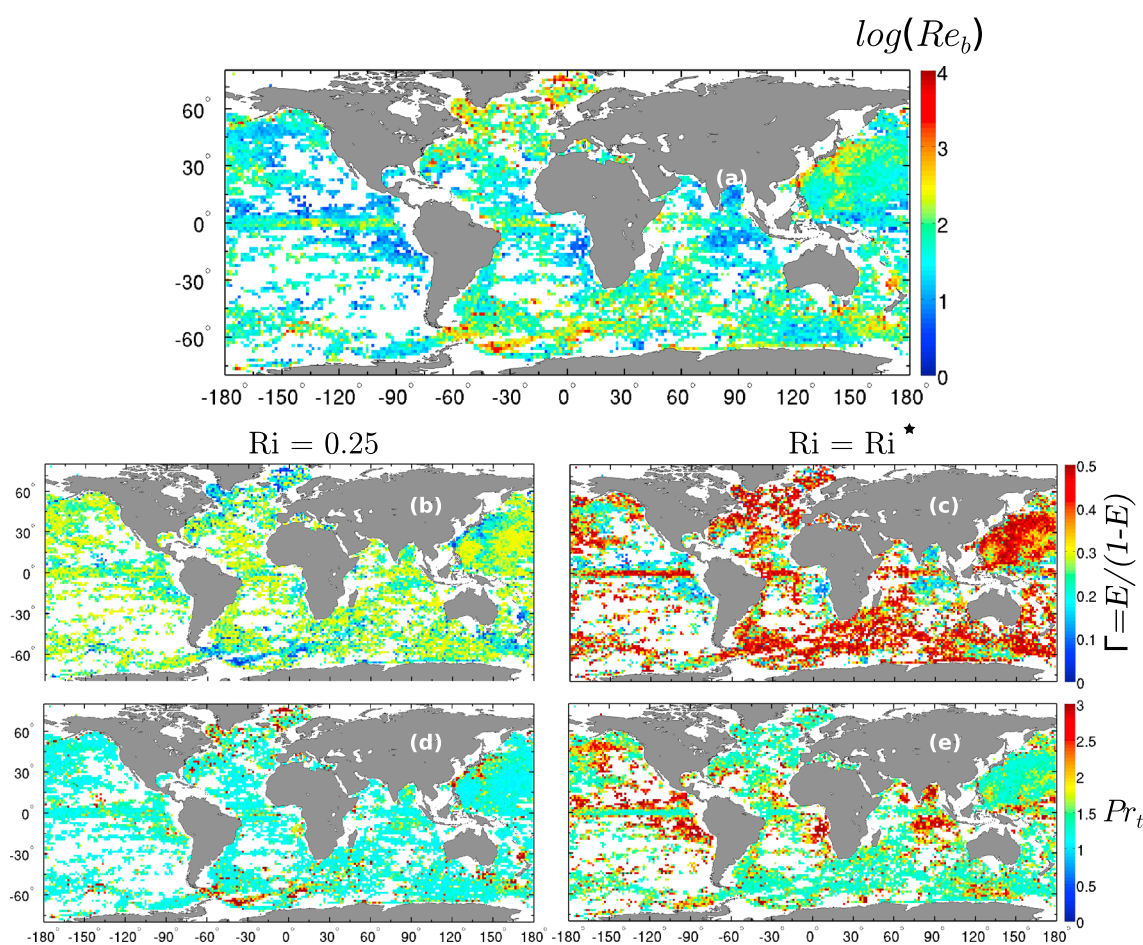
While the first notion was suggested by *Smyth et al.* [2001], the second assumption follows from the universality conjecture of SP. The latter is based on the striking similarity observed between the DNS results of the KH-ansatz and those of a homogeneously sheared and stratified flow [*Shih et al.*, 2005], despite their different forcing mechanism and spatial and temporal scales.

Under these assumptions and according to the earlier discussions, global maps of  $E$  (or  $\Gamma$ ) and  $Pr_t$  require a priori knowledge of the global distribution of both  $Ri$  and  $Re_b$ . While microstructure measurements may provide the needed information, their global distribution does not provide adequate spatial coverage [*Waterhouse et al.*, 2014]. Alternatively, the global array of Argo float profiles provides a suitable means of inferring the global distribution of  $Re_b$  by invoking the fine-scale parameterization approach to estimate turbulent dissipation [*Polzin et al.*, 2014] but lacks the measurements of the mean flow shear and hence  $Ri$  (see the supporting information for details).

While the assumptions underlying the fine-scale parameterization imply that it is inappropriate for some environments, it agrees in an unbiased manner with microstructure measurements within a factor of 2–3 in a variety of open ocean conditions including one region close to the equator [*Whalen et al.*, 2015]. Thus, the fine-scale parameterization is assumed to be appropriate for inferring the global-scale patterns in the following discussion despite its essential caveats as documented in *Whalen et al.* [2015].

## 3. Results

Global maps showing the 8 year mean values of the inferred  $Re_b = \epsilon/(\nu N^2)$  (assuming  $\nu = 10^{-6} \text{ m}^2/\text{s}$ ) as well as the parameterized estimates of the flux coefficient  $\Gamma = E/(1-E)$  (see (4)), and the turbulent Prandtl number,  $Pr_t = Ri_*/E$  are presented in Figure 2 for a depth interval of 250–450 m. The results of our analysis at all depth levels are consolidated in Figure 3 in terms of histograms of  $Re_b$ ,  $\Gamma$ ,  $K_p$ , and  $K_m$ , where the latter two have been computed using (2) and (3). The available depth-dependent inferred dissipation rates are associated with segments centered in 200 m bins covering the depth of 250–2000 m and are averaged into  $1.5^\circ$  square bins using data from *Whalen et al.* [2015]. Notice that globally constant values of either  $Ri = 0.25$  (a commonly

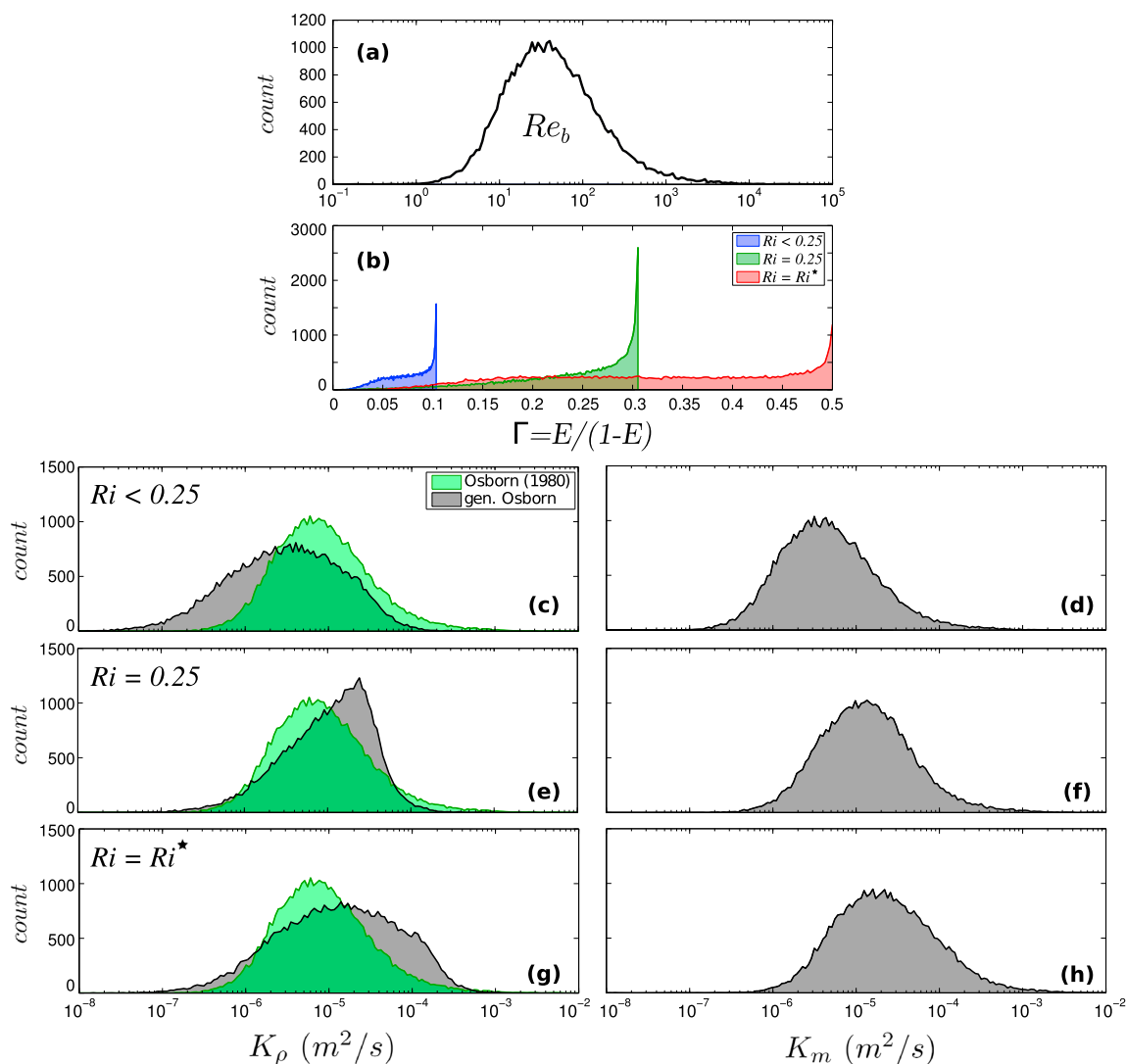


**Figure 2.** (a)  $Re_b = \epsilon / (\nu N^2)$  in which  $\epsilon$  and  $N^2$  are inferred from over 8 years (2006–2014) of Argo data. Estimates from high-vertical resolution data centered between 250 and 450 m are averaged over  $1.5^\circ$  square bins and plotted if they contain more than three dissipation rate estimates [Whalen et al., 2012, 2015]. (b, c) Parameterized global estimates of  $\Gamma = E / (1 - E)$ , as per (4) using the  $Re_b$  map in Figure 2a and assuming  $Ri = 0.25$  or  $Ri = Ri^*$ . (d, e) Parameterized global estimates of  $Pr_t = Ri^* / E$ , using the  $Re_b$  map in Figure 2a and assuming  $Ri = 0.25$  or  $Ri = Ri^*$ .

observed value associated with marginal instability, see Smyth and Moum [2013]) or  $Ri = Ri^* \sim 0.35 - 0.45$  are assumed in Figure 2. For illustrative purposes, Figure 3 also assumes  $Ri = 0.1 < 0.25$  in addition to the previous two  $Ri$  values. As noted earlier,  $Ri^*$  corresponds to the highest  $E^*$  point in the parameter space of Figure 1a. The precise value of  $Ri^*$  depends on parameterization details and will not affect our discussion to follow.

The results in Figures 2a and 3a demonstrate that  $Re_b$  varies over 4 orders of magnitude, ranging between  $\mathcal{O}(10^0)$  and  $\mathcal{O}(10^4)$ . The horizontally averaged value of  $Re_b$  lies in the range 90 – 180 for all depth levels. Commensurate variations of  $E$  and  $Pr_t$  are therefore expected based on the mixing properties of small-scale shear-driven stratified turbulence. In particular as shown in Figures 2b, 2c, and 3b, for  $Ri = 0.25$  (or  $Ri = Ri^*$ ), the value of  $\Gamma = E / (1 - E)$  varies in the range  $0 \leq \Gamma \leq 0.25$  (or  $0 \leq \Gamma \leq 0.5$  for  $Ri = Ri^*$ ) throughout the ocean with a global mean of  $\Gamma \sim 0.25$  (or  $\Gamma \sim 0.3$  for  $Ri = Ri^*$ ). These estimates may be reduced by a factor of 1.2–1.5 if the role of  $Pr$  is considered [see Salehipour et al., 2015, Figure 14d].

Depending on  $Ri$ , there may also exist prominent variability in regions with either high or low  $Re_b$ . For example, Figure 2b at  $Ri = 0.25$  would suggest that in the Southern Ocean  $\Gamma$  is significantly less than the global mean value of  $\sim 0.2$  in the Drake Passage, as well as over the Southwest Indian and Macquarie Ridges. This reduced mixing efficiency is due to the intensified turbulence leading to high values of  $Re_b$  in these regions which are characterized by rough bottom topography. The pattern of energetic turbulence and hence reduced mixing efficiency over regions of rough topography in the Southern Ocean seems to persist throughout the entire top 2 km of the ocean assuming that  $Ri = 0.25$ . North of New Zealand, mixing also seems to be relatively



**Figure 3.** Histograms of (a)  $Re_b = \epsilon/(\nu N^2)$  and (b)  $\Gamma = E/(1 - E)$  based on the new parameterization with constant values of  $Ri$ ; (c, e, and g)  $K_\rho$  based on the Osborn and generalized-Osborn formulae; (d, f, and h)  $K_m$  as per (3). Argo data at all depth levels are employed in constructing these histograms.

inefficient due to high estimates of  $Re_b$ . In contrast, examples of much lower estimates of mixing efficiency that are due to extremely low values of  $Re_b$  include the Bay of Bengal, the Peru Basin, and the Angola Plain.

Figures 2d and 2e also depict the global distribution of the parameterized  $Pr_t$  for the same two  $Ri$  values. Our results do not support the use of  $Pr_t = 10$  [Large et al., 1994; Danabasoglu et al., 2012] unless the ocean is assumed to be quiescent (i.e.,  $Re_b = \mathcal{O}(10^0)$ ) or extremely energetic (i.e.,  $Re_b \gtrsim \mathcal{O}(10^4)$ ). For  $Ri = 0.25$ ,  $Pr_t$  reaches 3 sporadically (e.g., in the Drake Passage) but it is generally lower in all depth levels. The globally averaged estimates of our parameterized  $Pr_t$  are much closer to  $Pr_t \approx 1 - 2$  for the top 2 km of the ocean. These global estimates may increase by a factor of 1.2–1.5 if the role of  $Pr$  is considered. The sensitivity of ocean general circulation models (OGCMs) to such changes in  $Pr_t$  remains unknown.

The histograms in Figures 3c–3h demonstrate the direct implications of the proposed parameterizations upon the characterization of  $K_\rho$  and  $K_m$ . As expected,  $K_m$  is insensitive to the variability of  $E$  (because  $E \rightarrow 1$ ) but changes with  $Ri$  as  $K_m/\nu = RiRe_b/(1 - E)$  (see (3)). On the other hand, the distribution of  $K_\rho$  is highly sensitive to  $E$  and differs substantially with that estimated by the Osborn method based on the canonical spatially invariant fixed value of  $\gamma = 0.2$  (see (1)).

If all small-scale mixing events were assumed to be optimally efficient (i.e.,  $Ri \sim Ri^*$ , see Figure 3g), mixing would typically occur more efficiently than is represented by  $\gamma = 0.2$ , causing a larger  $K_\rho$  than is predicted

by the Osborn formulation. On the other hand, if  $Ri \ll Ri^*$  or  $Ri \gg Ri^*$  (e.g., see Figure 3c), the flux coefficient and therefore  $K_p$  would be smaller than the values that are produced using the Osborn formulation (e.g., in the Southern Ocean and in the vicinity of the Drake passage).

The Diapycnal and Isopycnal Mixing Experiment in the Southern Ocean have focused on turbulent mixing through both microstructure measurements [St. Laurent *et al.*, 2012] and tracer release experiments [Ledwell *et al.*, 2011; Watson *et al.*, 2013]. The former method employs the Osborn formula with  $\gamma = 0.2$ . The average estimates of  $K_p$  inferred from these two methods differ slightly. While the difference between the measurement seasonality has been suggested to be an underlying reason for this discrepancy, the essential dependence of  $K_p$  on a variable mixing efficiency, as suggested by the histograms in Figure 3, may provide another, perhaps complementary, explanation.

#### 4. Conclusion

Based on our improved understanding of small-scale shear-induced stratified turbulence produced through Kelvin-Helmholtz instability (i.e., the KH-ansatz), we have demonstrated in this letter the variability of mixing efficiency ( $E$ ) and turbulent Prandtl number ( $Pr_t$ ) that arises because of their dependence upon both the buoyancy Reynolds number,  $Re_b = \epsilon/(\nu N^2)$ , as well as the gradient Richardson number,  $Ri = N^2/S^2$ . New multiparameter parameterizations for both  $E$  and  $Pr_t$  have been introduced which directly imply that  $K_p$  and  $K_m = Pr_t K_p$  should incorporate the influence of not only spatially varying turbulent dissipation ( $\epsilon$ ) and density stratification ( $N^2$ ) but also a spatially varying  $E$  and  $Pr_t$ .

By inferring  $Re_b$  from Argo float data and by assuming globally constant  $Ri$  values, we have shown that the resulting global maps of  $\Gamma = E/(1-E)$  (i) depend critically on  $Ri$  and (ii) may involve significant spatial variability which contrasts substantially with the assumption of a canonical flux coefficient of 0.2. Furthermore, the nonmonotonic dependence of  $E$  on  $Re_b$  at moderate levels of stratification implies that  $K_m \gg K_p$ , in contrast to the common assumption that  $Pr_t = 1$  for highly turbulent flows. Nevertheless, such high  $Pr_t$  values are only sporadically observed in the top 2 km of the ocean which is why on average  $Pr_t \approx 1 - 2$ .

These interesting findings should be further examined for deep and abyssal waters where the diapycnal mixing structure is understood to be strongly controlled by dissipation of the internal tide forced either by the flow of the barotropic tide [Garrett and Kunze, 2007; Klymak *et al.*, 2012] or geostrophic eddies [Nikurashin and Ferrari, 2011; Nikurashin *et al.*, 2013] over ocean bottom topography.

The currently prevalent view regarding the magnitude and depth-dependent trend of  $K_p$  [see, e.g., Waterhouse *et al.*, 2014] is entirely based on the conventional approach of utilizing the Osborn formula with a flux coefficient of 0.2. Our revised parameterization for mixing efficiency embodied in the generalized-Osborn formula suggests that reconsideration of this view is warranted.

Previous studies have shown significant sensitivity of large-scale climate models to the horizontal and vertical distribution of  $K_p$  [Jochum, 2009; Jayne, 2009; Saenko *et al.*, 2012; Melet *et al.*, 2013]. For example, in the context of paleoclimate modeling, Peltier and Vettoretti [2014] and Vettoretti and Peltier [2015] have demonstrated that the intensely nonlinear Dansgaard-Oeschger oscillation, that is characteristic of cold glacial climate conditions, is highly sensitive to the representation of  $K_p$ . This letter provides the physical rationale for constraining the variability of  $E$  and subsequently for developing a revised representation of  $K_p$  to be employed in future sensitivity studies of both the modern and the glacial ocean.

Although the proposed parameterizations significantly extend the previous single-parameter representations (e.g., that of Shih *et al.* [2005] for  $E$ ) to a multiparameter framework as previously suggested to be necessary by Mater and Venayagamoorthy [2014], they should be regarded as a preliminary attempt to connect the idealized studies of individual stratified mixing events to our understanding of global ocean mixing. Furthermore, the new characterizations may require further refinement prior to their implementation in an OGCM in order to relate the coarsely resolved OGCM estimates of  $Ri$  to those associated with small-scale turbulence as defined and employed in this letter.

Future efforts will be required (i) to understand the mixing properties of stratified turbulent flows induced by processes other than mean flow shear: e.g., by double-diffusion; (ii) to investigate whether there might be a “universal” characteristic behavior associated with stratified turbulent mixing as, e.g., conjectured by SP; (iii) to study the existence and strength of turbulent mixing in *strongly* stratified flows with  $Ri > 0.5$ ; and



(iv) to understand the role of molecular Prandtl number ( $Pr$ ) on high- $Re_b$  shear-dominated flows. This will enable future adjustments to this parameterization in order to better represent the thermal, saline, and mixed stratification regimes observed in the ocean.

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